

## Further ensemble averages [tln55]

Probability density in one-particle phase space:

$$\rho_l(\mathbf{q}, \mathbf{p}) = \langle \delta(\mathbf{q}_l - \mathbf{q}) \delta(\mathbf{p}_l - \mathbf{p}) \rangle.$$

Position distribution and momentum distribution:

$$\rho_l(\mathbf{q}) = \langle \delta(\mathbf{q}_l - \mathbf{q}) \rangle, \quad \rho_l(\mathbf{p}) = \langle \delta(\mathbf{p}_l - \mathbf{p}) \rangle.$$

Distribution of distances and relative momenta between pairs of particles:

$$f_{lm}(r) = \langle \delta(r - |\mathbf{q}_l - \mathbf{q}_m|) \rangle, \quad F_{lm}(P) = \langle \delta(P - |\mathbf{p}_l - \mathbf{p}_m|) \rangle.$$

Average distance between pairs of particles:

$$\langle r_{lm} \rangle = \langle |\mathbf{q}_l - \mathbf{q}_m| \rangle = \int_0^\infty dr r f_{lm}(r).$$

Average magnitude of relative momentum between pairs of particles:

$$\langle P_{lm} \rangle = \langle |\mathbf{p}_l - \mathbf{p}_m| \rangle = \int_0^\infty dP P F_{lm}(P).$$

**Applications to the classical ideal gas:**

$$\text{Noninteracting particles: } h(\mathbf{q}, \mathbf{p}) = \frac{p^2}{2m} \Rightarrow \tilde{Z} = \frac{V}{\lambda_T^3}, \quad \lambda_T = \sqrt{\frac{h^2}{2\pi m k_B T}}.$$

$$\Rightarrow \rho_l(\mathbf{q}, \mathbf{p}) = V^{-1} (2\pi m k_B T)^{-3/2} e^{-p^2/2m k_B T}.$$

$$\Rightarrow \rho_l(\mathbf{q}) = \int d^3p \rho_l(\mathbf{q}, \mathbf{p}) = \frac{1}{V}.$$

$$\Rightarrow \rho_l(\mathbf{p}) = \int d^3q \rho_l(\mathbf{q}, \mathbf{p}) = (2\pi m k_B T)^{-3/2} e^{-p^2/2m k_B T}.$$

$$\rho_l(\mathbf{p}) d^3p = f(\mathbf{v}) d^3v \Rightarrow f(\mathbf{v}) = \left( \frac{m}{2\pi k_B T} \right)^{-3/2} e^{-mv^2/2k_B T}.$$

The spatial distribution of ideal gas particles in a uniform gravitational field (law of atmospheres) is calculated in exercise [tex79].

The distribution of relative momenta between pairs of ideal gas particles is calculated in exercise [tex80].