



Self-Gravitating Lattice Gas

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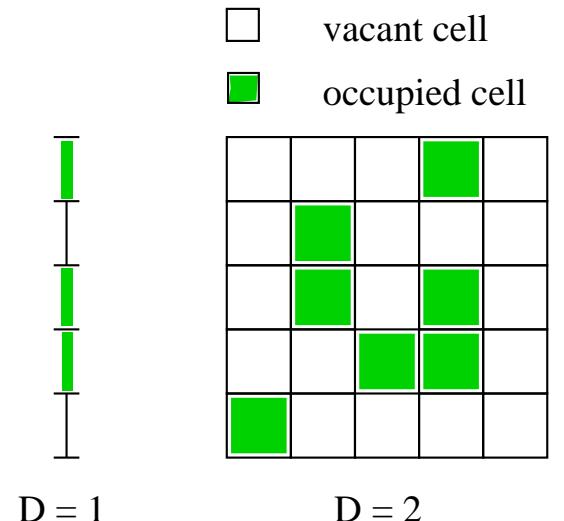
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Ideal Lattice Gas

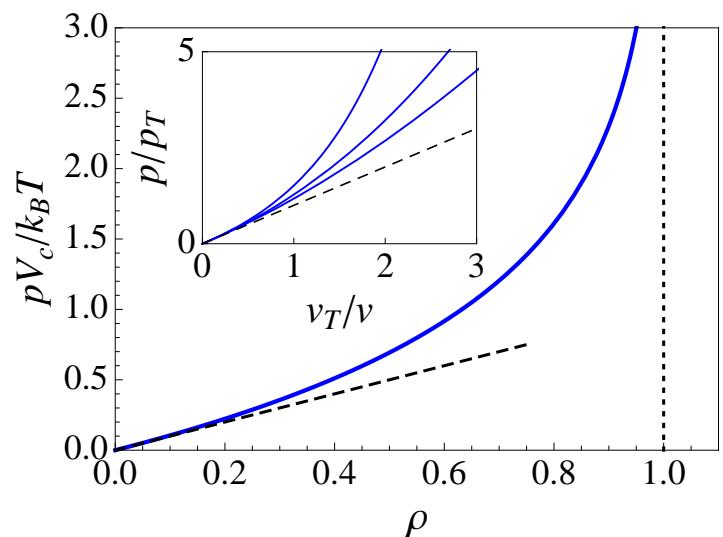


- volume of cell: V_c
- number of cells: N_c
- number of particles: N
- density: $\rho = N/N_c$



Equations of state (EOS):

- Lattice gas: $\frac{pV_c}{k_B T} = -\ln(1-\rho)$
 - FD gas: $\frac{p}{P_T} = f_{D/2+1}(z), \quad \rho \doteq \frac{v_T}{v} = f_{D/2}(z)$
- $$p \sim \rho^{(D+2)/D} \quad (\text{isotherm})$$



Equilibrium Condition



Heterogeneous environment:

- Presence of external potential $\mathcal{U}(\mathbf{r})$.
- Equilibrium condition has two parts:
 - *equation of state* (EOS):
(local) thermal equilibrium from minimized free energy density,
 - *equation of motion* (EOM):
mechanical equilibrium from balanced forces.
- Equilibrium establishes density profile $\rho(\mathbf{r})$ and pressure profile $p(\mathbf{r})$.
- Temperature T remains uniform.
- Potential $\mathcal{U}(\mathbf{r})$ here represented by long-range gravitational interactions.
- Mean-field approach proven exact.
- Ensemble inequivalence to be heeded.

Gravity in $\mathcal{D} = 1, 2, 3$ Dimensions



Interaction potential \mathcal{V}_{ij} and force F_{ij} between occupied cells of mass m_c at distance r_{ij} :

$$\frac{\mathcal{V}_{ij}}{Gm_c^2} \begin{cases} r_{ij} & : \mathcal{D} = 1, \\ \ln r_{ij} & : \mathcal{D} = 2, \\ -r_{ij}^{-1} & : \mathcal{D} = 3. \end{cases} \quad F_{ij} = -\frac{Gm_c^2}{r_{ij}^{\mathcal{D}-1}}.$$

Gauss' law: $\oint d\mathbf{A} \cdot \mathbf{g}(\mathbf{r}) = -\mathcal{A}_{\mathcal{D}} G m_{\text{in}}, \quad \mathcal{A}_{\mathcal{D}} = \frac{2\pi^{\mathcal{D}/2}}{\Gamma(\mathcal{D}/2)} = \begin{cases} 2 & : \mathcal{D} = 1, \\ 2\pi & : \mathcal{D} = 2, \\ 4\pi & : \mathcal{D} = 3. \end{cases}$

Symmetric self-gravitating cluster with center of mass at $\mathbf{r} = 0$.

Gravitational field: $\mathbf{g}(r) = -\frac{m_{\text{in}}}{r^{\mathcal{D}-1}}.$

Gravitational potential $\mathcal{U}(r)$ from $\mathbf{g}(r) = -\frac{d\mathcal{U}}{dr}.$

Self-Gravitating Lattice Gas at Equilibrium



Thermal and hydrostatic equilibrium:

- EOS: $\frac{p(r)V_c}{k_B T} = -\ln(1 - \rho(r))$,
- EOM: $\frac{dp}{dr} = \frac{m_c}{V_c} \rho(r) g(r)$, $g(r) = -\frac{G m_{\text{in}}}{r^{\mathcal{D}-1}}$, $m_{\text{in}} = \frac{m_c}{V_c} \int_0^r dr' (\mathcal{A}_{\mathcal{D}} r'^{\mathcal{D}-1}) \rho(r')$.

Scaled variables:

- $\hat{r} \doteq \frac{r}{r_s}$, $\hat{p} \doteq \frac{p}{p_s}$, $\hat{T} \doteq \frac{k_B T}{p_s V_c}$, $\hat{\mathcal{U}} \doteq \frac{\mathcal{U}}{p_s V_c}$,
- $r_s^{\mathcal{D}} = \frac{N V_c \mathcal{D}}{\mathcal{A}_{\mathcal{D}}}$, $p_s = \frac{\mathcal{A}_{\mathcal{D}} G}{2\mathcal{D}} \frac{m_c^2}{V_c^2} r_s^2$.

Relations between profiles:

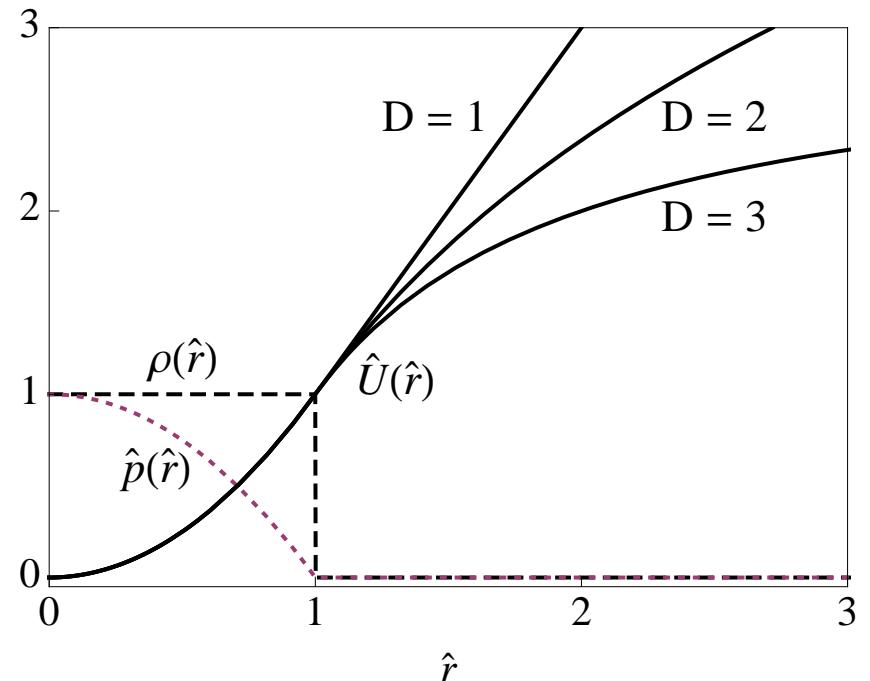
- EOS & EOM: $\hat{p}(\hat{r}) = -\hat{T} \ln(1 - \rho(\hat{r}))$, $\frac{d\hat{p}}{d\hat{r}} = -2\mathcal{D} \rho(\hat{r}) \int_0^{\hat{r}} d\hat{r}' \left(\frac{\hat{r}'}{\hat{r}}\right)^{\mathcal{D}-1} \rho(\hat{r}')$,
- $\hat{U}(\hat{r}) = \hat{T} \ln \left(\frac{1 - \rho(\hat{r})}{1 - \rho(0)} \frac{\rho(0)}{\rho(\hat{r})} \right)$. Reference value: $\hat{U}(0) = 0$.

Profiles at $T = 0$: Solid Cluster



- $\rho(\hat{r}) = \begin{cases} 1 & : 0 \leq \hat{r} < 1, \\ 0 & : \hat{r} > 1, \end{cases}$
- $\hat{p}(\hat{r}) = \begin{cases} 1 - \hat{r}^2 & : 0 \leq \hat{r} < 1, \\ 0 & : \hat{r} > 1, \end{cases}$
- $\hat{\mathcal{U}}(\hat{r}) = \hat{r}^2 \quad : 0 \leq \hat{r} \leq 1,$

$$\hat{\mathcal{U}}(\hat{r}) = \begin{cases} 2\hat{r} - 1 & : \mathcal{D} = 1, \\ 2 \ln \hat{r} + 1 & : \mathcal{D} = 2, \\ 3 - 2/\hat{r} & : \mathcal{D} = 3, \end{cases} \quad \hat{r} \geq 1.$$



Escape velocity finite in $\mathcal{D} > 2$.

Stable clusters of finite mass exist

- at all \hat{T} ($\mathcal{D} = 1$),
- at $\hat{T} < \hat{T}_c$ ($\mathcal{D} = 2$),
- at $\hat{T} = 0$ only ($\mathcal{D} = 3$).

Differential Equations for Profiles



Analysis of EOS & EOM at $\hat{T} > 0$:

- density: $\rho'' + \frac{\mathcal{D}-1}{\hat{r}}\rho' - \frac{1-2\rho}{\rho(1-\rho)}(\rho')^2 + \frac{2\mathcal{D}}{\hat{T}}\rho^2(1-\rho) = 0, \quad \rho'(0) = 0,$
- pressure: $\hat{p}'' + \frac{\mathcal{D}-1}{\hat{r}}\hat{p}' - \frac{1-\rho}{\hat{T}\rho}(\hat{p}')^2 + 2\mathcal{D}(\rho)^2 = 0, \quad \hat{p}'(0) = 0,$
- potential: $\hat{\mathcal{U}}'' + \frac{\mathcal{D}-1}{\hat{r}}\hat{\mathcal{U}}' - 2\mathcal{D}\rho = 0, \quad \hat{\mathcal{U}}(0) = \hat{\mathcal{U}}'(0) = 0,$

with $\rho(\hat{r}) = 1 - e^{-\hat{p}(\hat{r})/\hat{T}}$; $\rho(\hat{r}) = \frac{1}{e^{(\hat{\mathcal{U}}(\hat{r})-\hat{\mu})/\hat{T}} + 1}$, $\hat{\mu} \doteq \frac{\mu}{p_s V_c} = -\hat{T} \ln \left(\frac{1 - \rho(0)}{\rho(0)} \right)$.

Additional boundary conditions:

- density: $\mathcal{D} \int_0^\infty d\hat{r} \hat{r}^{\mathcal{D}-1} \rho(\hat{r}) = 1,$
- pressure: $2\mathcal{D}(\mathcal{D}-1) \int_0^\infty d\hat{r} \hat{r}^{2\mathcal{D}-3} \hat{p}(\hat{r}) = 1,$
- in $\mathcal{D} = 1$ only: $\hat{p}(0) = 1, \quad \rho(0) = 1 - e^{-1/\hat{T}}.$

Ideal Classical Gas Limit (1)



ICG assumption: low density $\rho(\hat{r}) \ll 1$ realized throughout the cluster.

ODE for density profile: $\frac{\rho''}{\rho} + \frac{\mathcal{D} - 1}{\hat{r}} \frac{\rho'}{\rho} - \left(\frac{\rho'}{\rho} \right)^2 + \frac{2\mathcal{D}}{\hat{T}} \rho = 0,$

with boundary conditions as in ILG.

Analytic solutions for finite clusters:

- $\mathcal{D} = 1$: $\rho(\hat{r})_{\text{ICG}} = \frac{1}{\hat{T}} \operatorname{sech}^2 \left(\frac{\hat{r}}{\hat{T}} \right),$

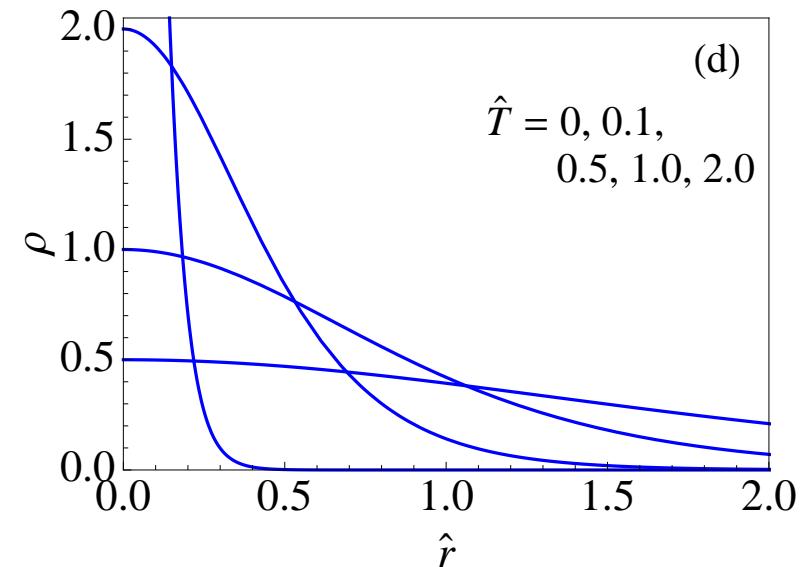
solution normalized at all \hat{T} ,

exponential decay: $\rho(\hat{r}) \sim e^{-2\hat{r}/\hat{T}}.$

- $\mathcal{D} = 2$: $\rho(\hat{r})_{\text{ICG}} = \frac{4c}{\hat{T}} \left[1 + 2c \left(\frac{\hat{r}}{\hat{T}} \right)^2 \right]^{-2},$

solution normalizable only for $\hat{T}_c = \frac{1}{2},$

power-law decay: $\rho(\hat{r}) \sim \hat{r}^{-4}.$



Ideal Classical Gas Limit (2)



ICG in $\mathcal{D} = 3$ confined to sphere of radius $\hat{R} \doteq R/r_s$.

Rescaled variables: $\bar{r} \doteq \hat{r}/\hat{R}$, $\bar{\rho} \doteq \hat{R}^3 \rho$, $\bar{T} \doteq \hat{R}\hat{T}$.

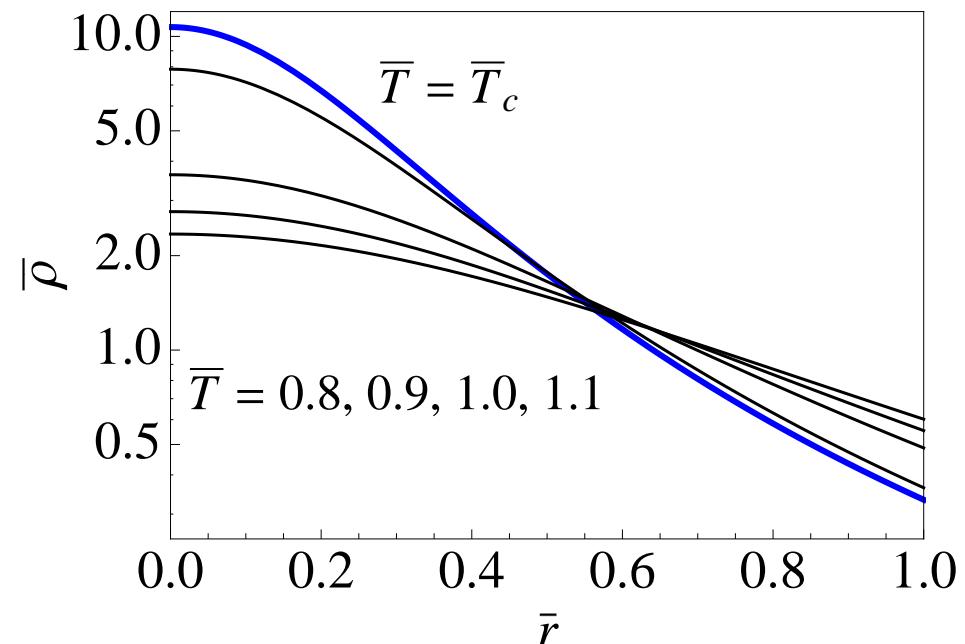
ODE: $\frac{\bar{\rho}''}{\bar{\rho}} + \frac{2}{\bar{r}} \frac{\bar{\rho}'}{\bar{\rho}} - \left(\frac{\bar{\rho}'}{\bar{\rho}} \right)^2 + \frac{6\bar{\rho}}{\bar{T}} = 0$, $\bar{\rho}'(0) = 0$, $3 \int_0^1 d\bar{r} \bar{r}^2 \bar{\rho}(\bar{r}) = 1$.

Normalizable solutions exist for $\bar{T} \geq \bar{T}_C = 0.794422\dots$

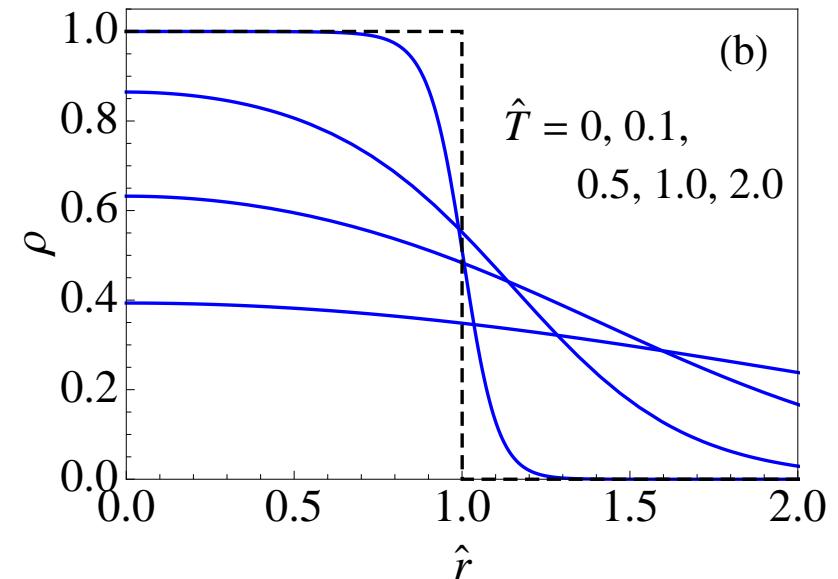
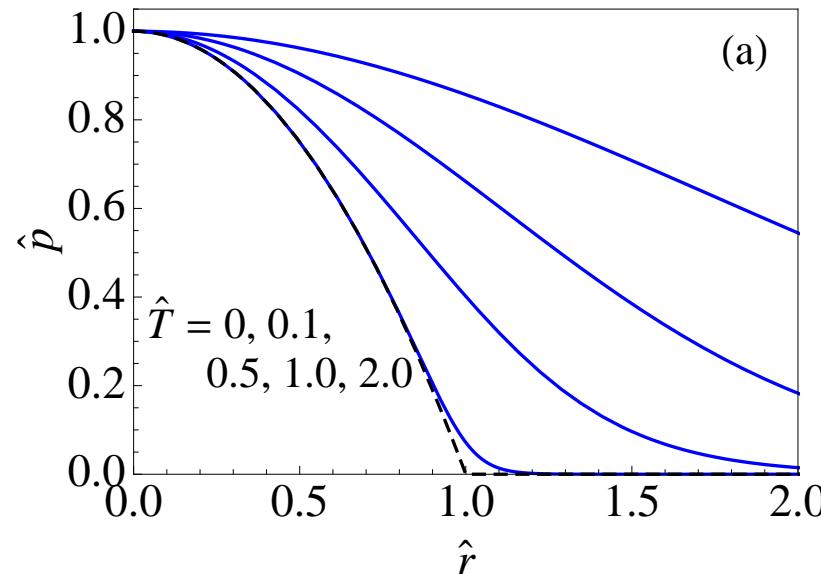
Vega and Sanchez (2006) predict

- spinodal point: $\bar{T}_C = 0.79442$,
- transition point: $\bar{T}_T = 0.8215$,

based on computer simulations.



ILG Profiles in $\mathcal{D} = 1$

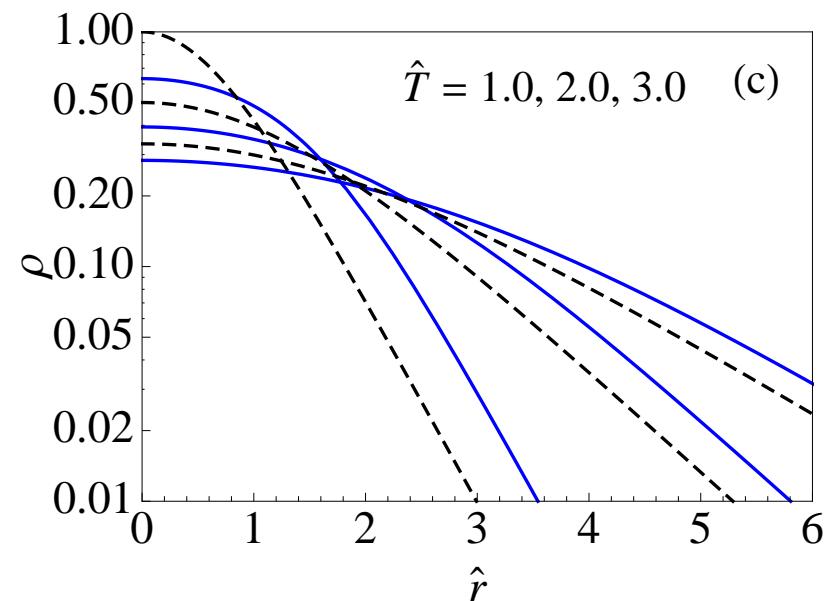


Analytic ICG profile:

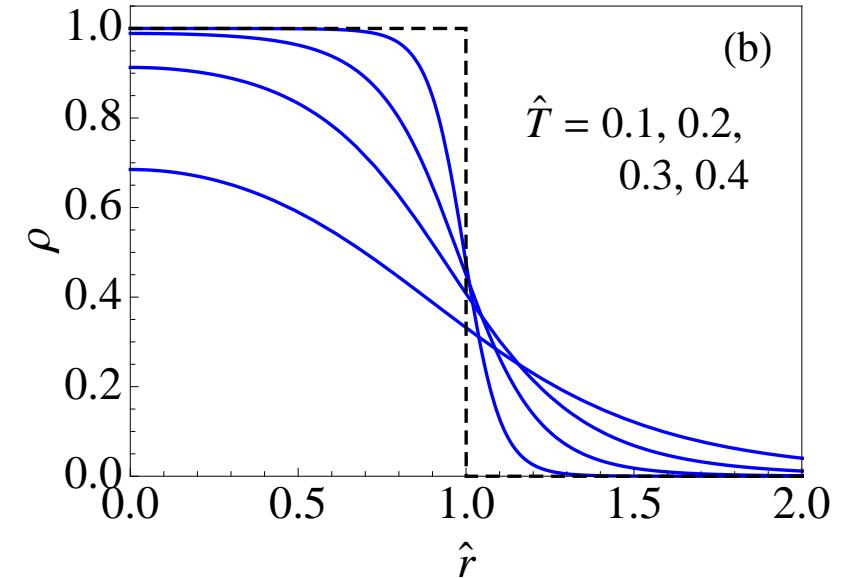
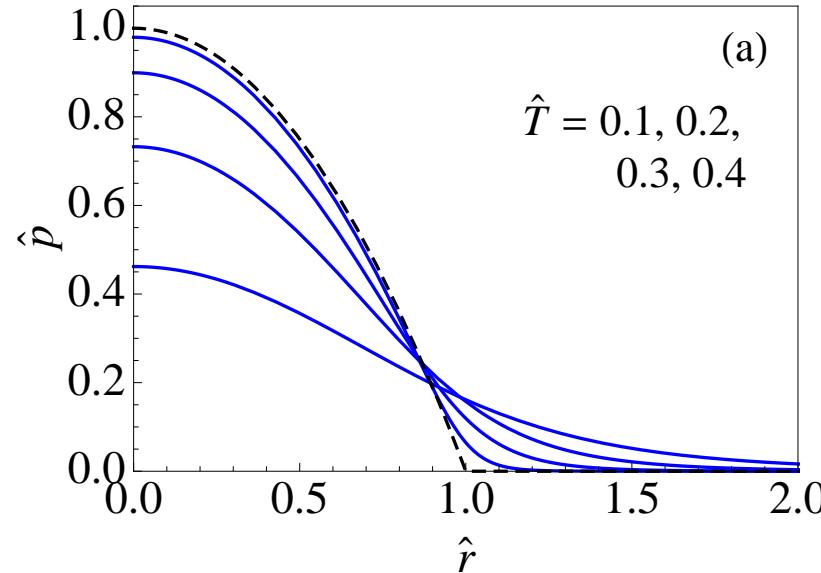
$$\rho(\hat{r})_{\text{ICG}} = \frac{1}{\hat{T}} \operatorname{sech}^2\left(\frac{\hat{r}}{\hat{T}}\right)$$

Exact ILG asymptotics:

$$\rho(\hat{r})_{\text{as}} \sim e^{-2\hat{r}/\hat{T}} \quad : \hat{r} \gg 1$$



ILG Profiles for $\mathcal{D} = 2$: Unconfined Space

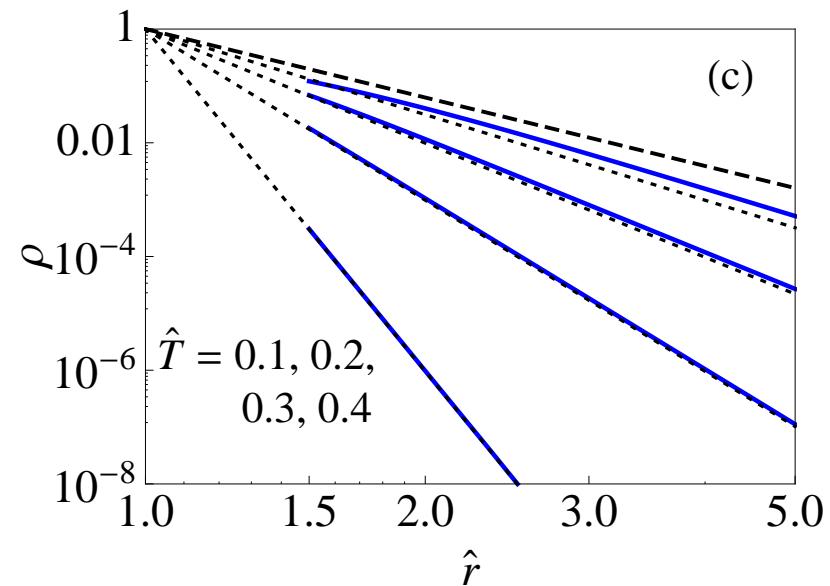


Critical ILG profile: $\hat{T} \lesssim \hat{T}_c = \frac{1}{2}$

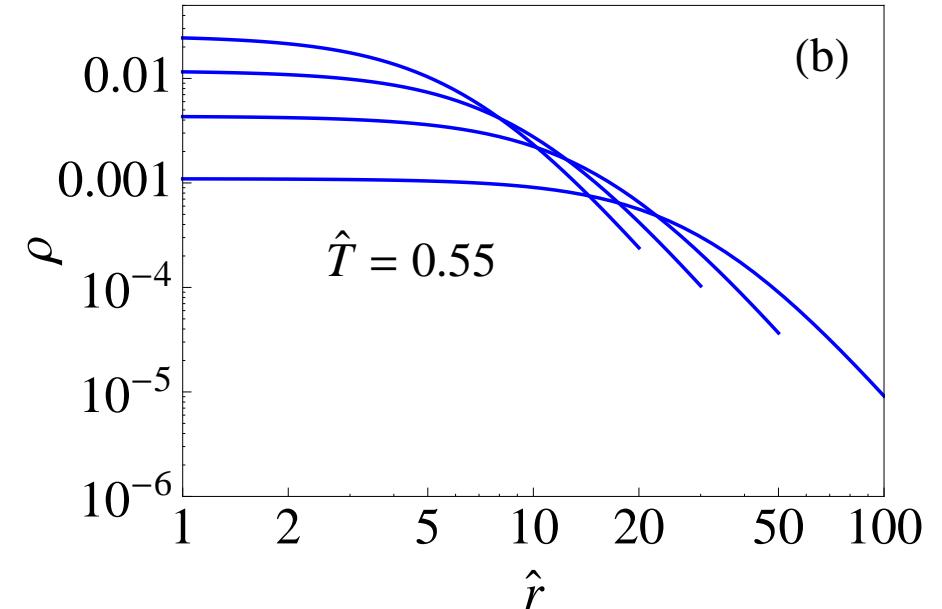
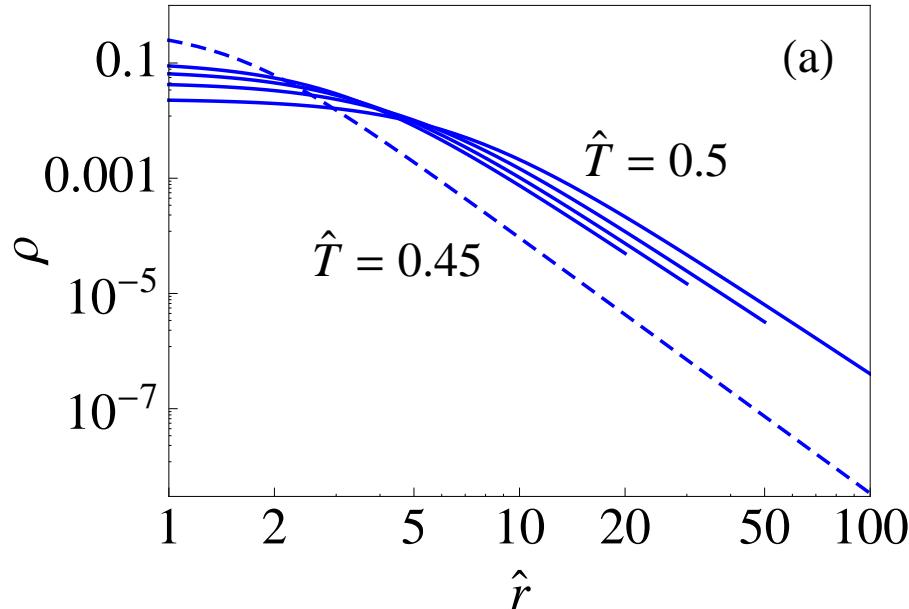
$$\rho(\hat{r})_{\text{ICG}} = \frac{4c}{\hat{T}} \left[1 + 2c \left(\frac{\hat{r}}{\hat{T}} \right)^2 \right]^{-2}$$

Exact ILG asymptotics: $0 \leq \hat{T} \leq \hat{T}_c$

$$\rho(\hat{r})_{\text{as}} \sim \hat{r}^{-2/\hat{T}} \quad : \quad \hat{r} \gg 1$$



ILG Profiles for $\mathcal{D} = 2$: Confined Space



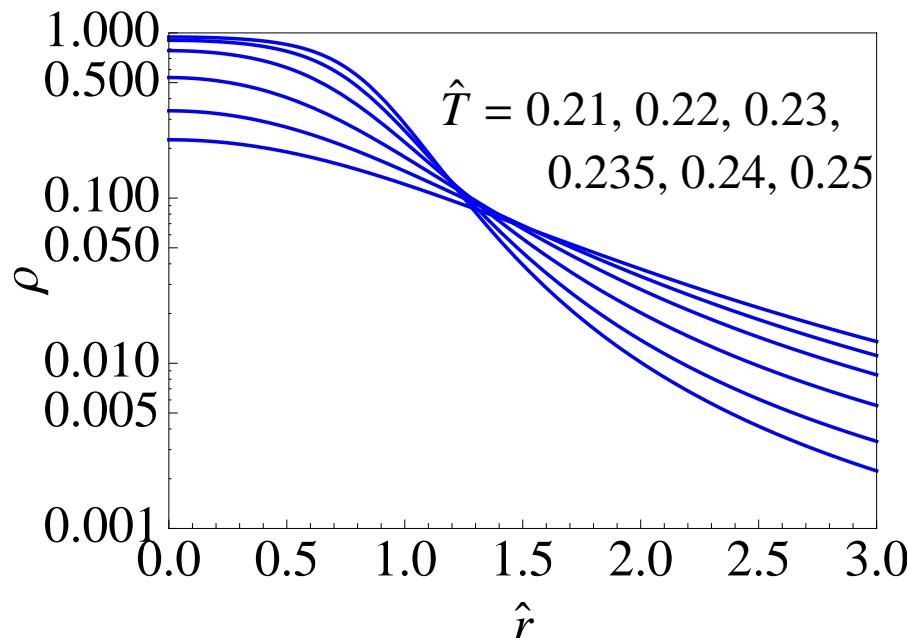
ILG in $\mathcal{D} = 2$ confined to disk of radius $\hat{R} = 20, 30, 50, 100$.

- $\hat{T} = 0.45$: stable cluster,
- $\hat{T} = 0.5$: stability limit,
- $\hat{T} = 0.55$: unstable cluster.

ILG Profiles for $\mathcal{D} = 3$: Confined Space

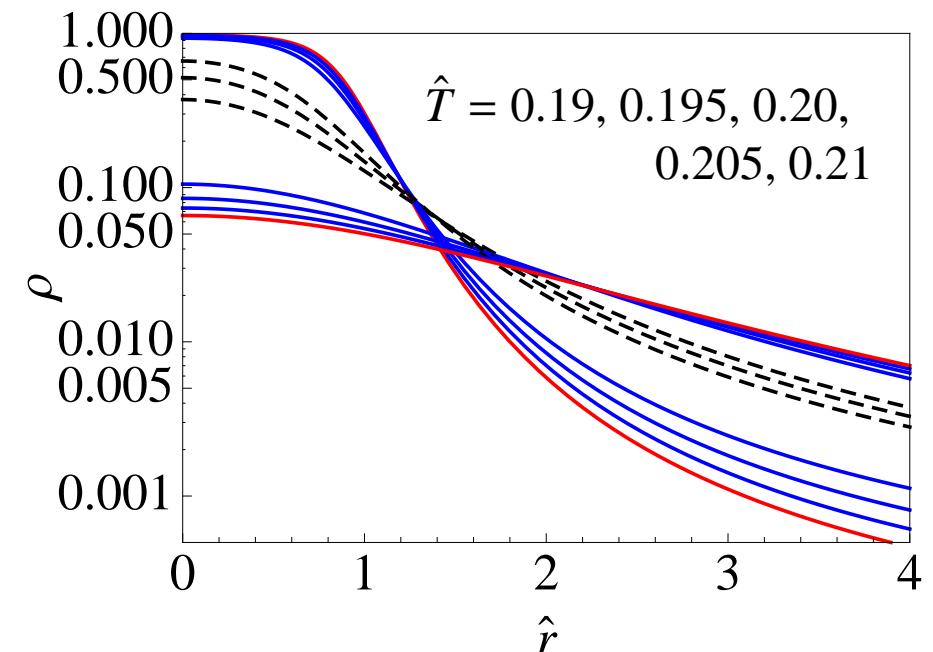


ILG in $\mathcal{D} = 3$ confined to sphere of radius \hat{R} .



$$\hat{R} = 3$$

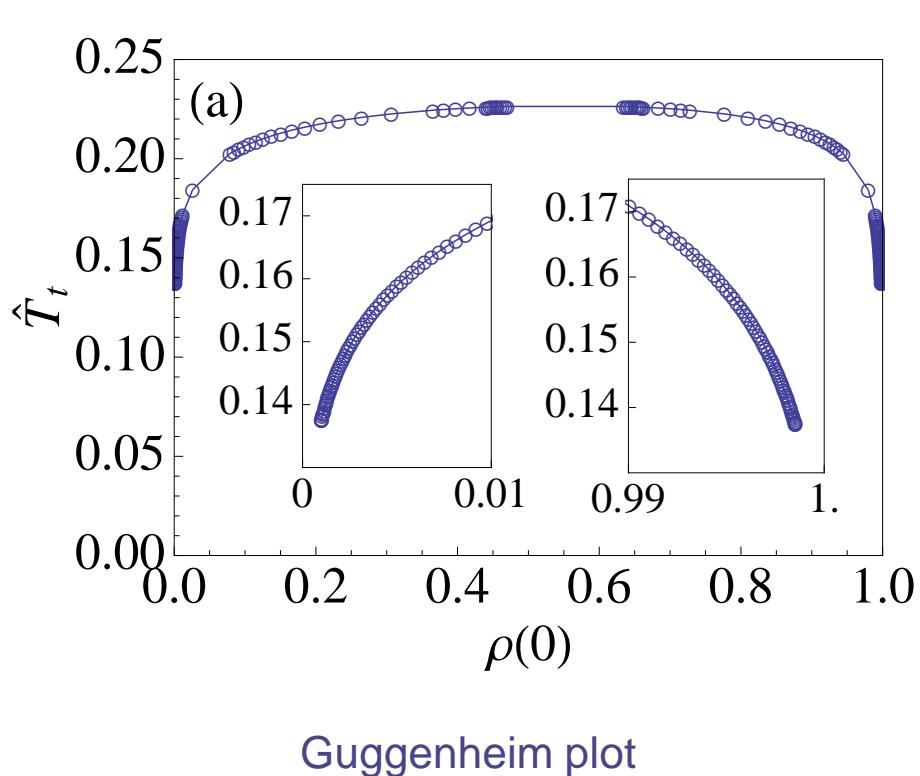
- one normalizable solution at all \hat{T} ,
- gradual cluster formation,
- no phase transition.



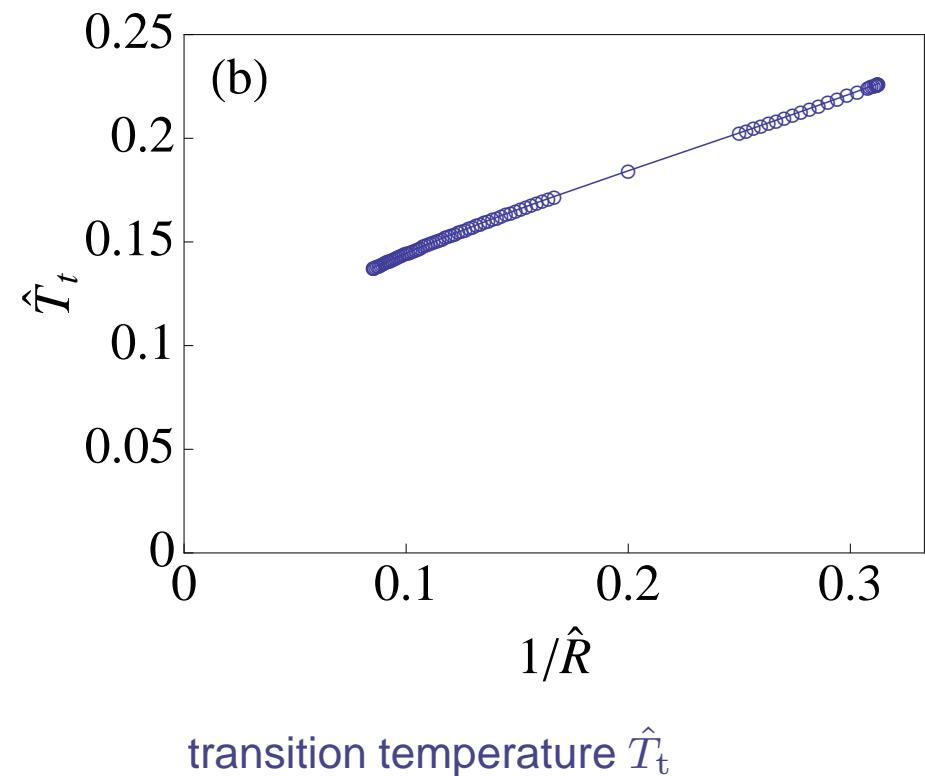
$$\hat{R} = 4$$

- three normalizable solutions in \hat{T} interval,
- two mechanically stable solutions,
- free energy minimum switches,
- first-order phase transition.

ILG Profiles for $\mathcal{D} = 3$: First-Order Transition



Guggenheim plot



transition temperature \hat{T}_t

- critical temperature: $\hat{T}_c \lesssim 0.23$,
- critical radius: $\hat{R}_c \gtrsim 3.1$

Open ILG Clusters



- Closed system: finite mass, confined or unconfined space.

$$\rho'' + \frac{\mathcal{D} - 1}{\hat{r}}\rho' - \frac{1 - 2\rho}{\rho(1 - \rho)}(\rho')^2 + \frac{2\mathcal{D}}{\hat{T}}\rho^2(1 - \rho) = 0, \quad \rho'(0) = 0, \quad \mathcal{D} \int_0^{\hat{R}} d\hat{r} \hat{r}^{\mathcal{D}-1} \rho(\hat{r}) = 1.$$

$$\hat{r} \doteq \frac{r}{r_s}, \quad \hat{p} \doteq \frac{p}{p_s}, \quad \hat{T} \doteq \frac{k_B T}{p_s V_c}, \quad \hat{\mathcal{U}} \doteq \frac{\mathcal{U} m_c}{p_s V_c}, \quad r_s^{\mathcal{D}} = \frac{N V_c \mathcal{D}}{\mathcal{A}_{\mathcal{D}}}, \quad p_s = \frac{\mathcal{A}_{\mathcal{D}} G}{2\mathcal{D}} \frac{m_c^2}{V_c^2} r_s^2.$$

- Open system: finite or infinite mass, unconfined space.

$$\tilde{\rho}'' + \frac{\mathcal{D} - 1}{\tilde{r}}\tilde{\rho}' - \frac{1 - 2\rho_0 \tilde{\rho}}{\tilde{\rho}(1 - \rho_0 \tilde{\rho})} (\tilde{\rho}')^2 + \tilde{\rho}^2(1 - \rho_0 \tilde{\rho}) = 0, \quad \tilde{\rho}(0) = 1, \quad \tilde{\rho}'(0) = 0.$$

$$\tilde{r} \doteq \sqrt{\frac{2\mathcal{D}\rho_0}{\hat{T}}} \hat{r}, \quad \tilde{\rho} \doteq \frac{\rho}{\rho_0}.$$

- ICG limit $\rho_0 \rightarrow 0$:

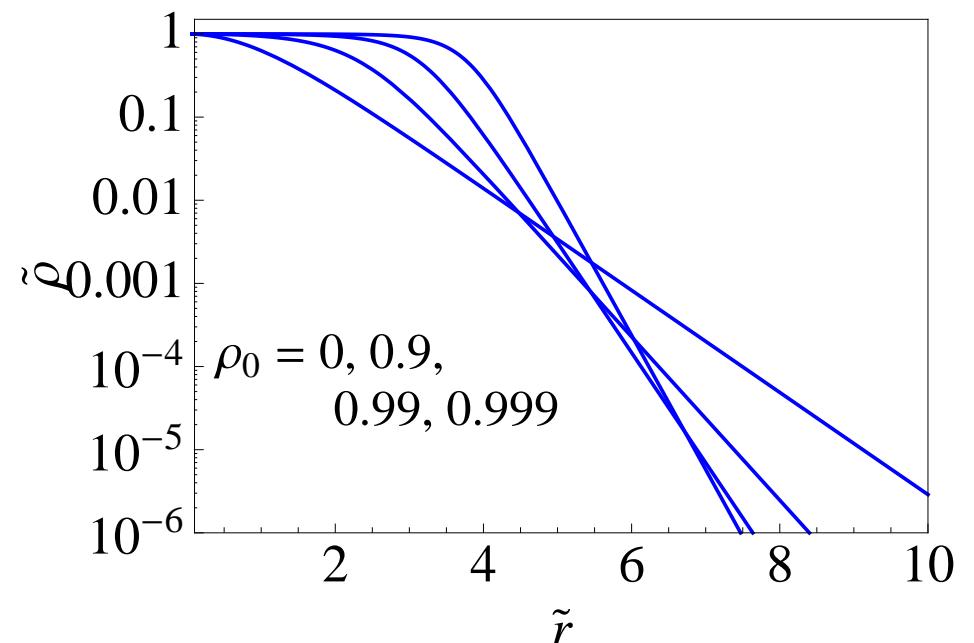
$$\frac{\tilde{\rho}''}{\tilde{\rho}} + \frac{\mathcal{D} - 1}{\tilde{r}} \frac{\tilde{\rho}'}{\tilde{\rho}} - \left(\frac{\tilde{\rho}'}{\tilde{\rho}} \right)^2 + \tilde{\rho} = 0 \quad \Rightarrow \quad \tilde{\rho}(\tilde{r}) \sim \begin{cases} e^{-\sqrt{2}\tilde{r}} & : \mathcal{D} = 1, \\ \tilde{r}^{-4} & : \mathcal{D} = 2, \\ \tilde{r}^{-2} & : \mathcal{D} = 3. \end{cases}$$

Open ILG Cluster in $\mathcal{D} = 1$



Use inverse function $\tilde{r}(\tilde{\rho})$ for analytic solution:

- $\tilde{r}'' + \frac{1-2\rho}{\rho(1-\rho)}\tilde{r}' - \frac{1}{\rho_0}\rho^2(1-\rho)(\tilde{r}')^3 = 0, \quad \rho = \rho_0\tilde{\rho}, \quad \tilde{r}(\rho_0) = 0, \quad \tilde{r}'(\rho_0) = -\infty$
- $\tilde{r}(\tilde{\rho}) = \sqrt{\frac{\rho_0}{2}} \int_{\tilde{\rho}}^1 \frac{d\tilde{\rho}'}{\tilde{\rho}'(1-\rho_0\tilde{\rho}')\sqrt{\ln \frac{1-\rho_0\tilde{\rho}'}{1-\rho_0}}} \quad : 0 \leq \tilde{\rho} \leq 1$
- $\tilde{\rho}(\tilde{r})_{\text{as}} \sim e^{-\nu_1(\rho_0)\tilde{r}}$
- $\nu_1(\rho_0) = \sqrt{-\frac{2}{\rho_0} \ln(1-\rho_0)}$
- $\tilde{\rho}(\tilde{r})_{\text{ICG}} = \operatorname{sech}^2 \left(\frac{\tilde{r}}{\sqrt{2}} \right)$
- finite mass, exponential decay law

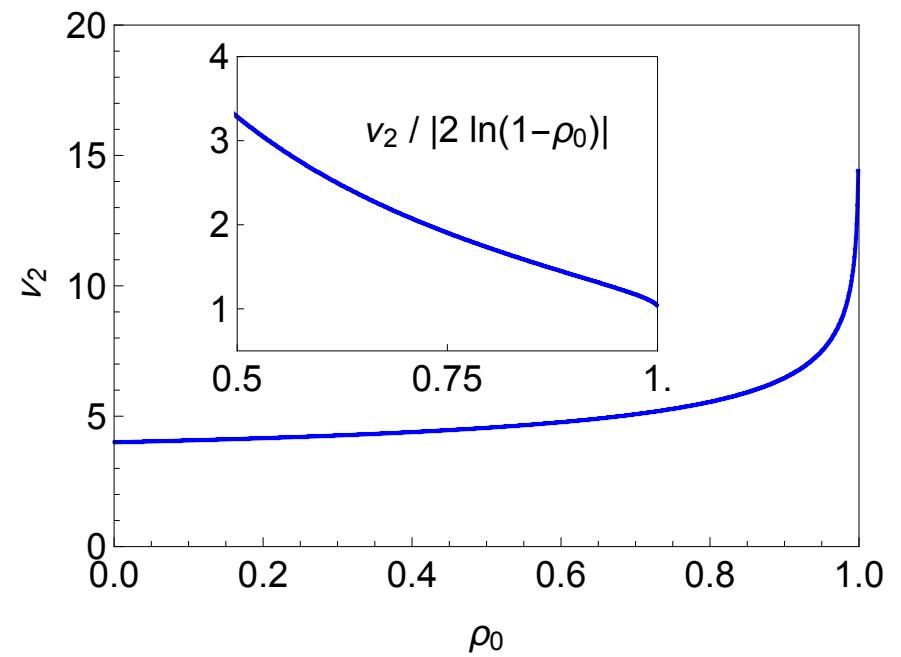
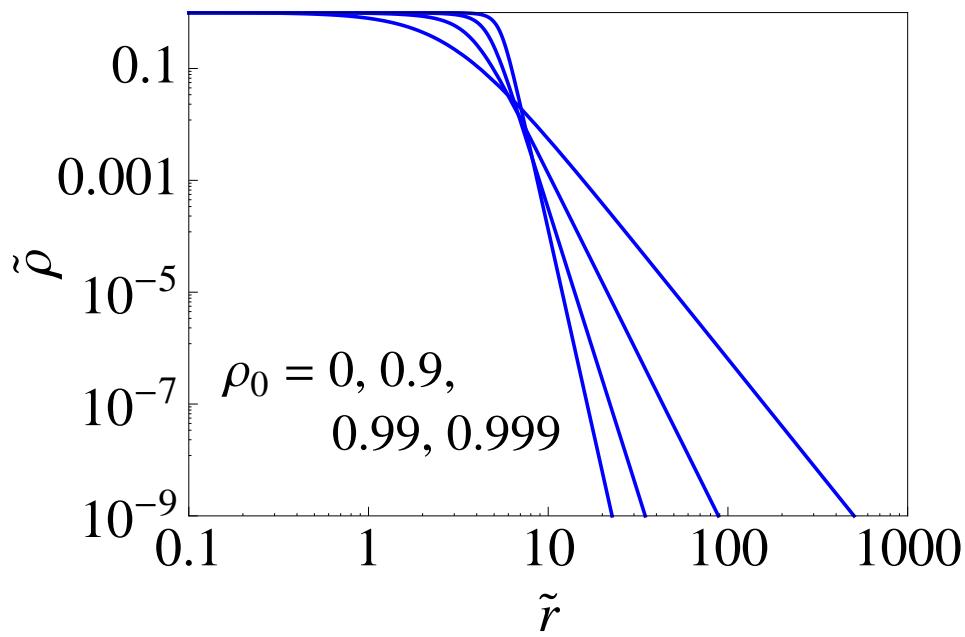


Open ILG Cluster in $\mathcal{D} = 2$



Numerical solution:

- $\tilde{\rho}(\tilde{r})_{\text{as}} \sim \tilde{r}^{-\nu_2(\rho_0)}$, $\nu_2(\rho_0) \xrightarrow{\rho_0 \rightarrow 1} -2 \ln(1 - \rho_0)$
- $\tilde{\rho}(\tilde{r})_{\text{ICG}} = \frac{1}{(1 + \tilde{r}^2/8)^2}$
- finite mass, algebraic decay law

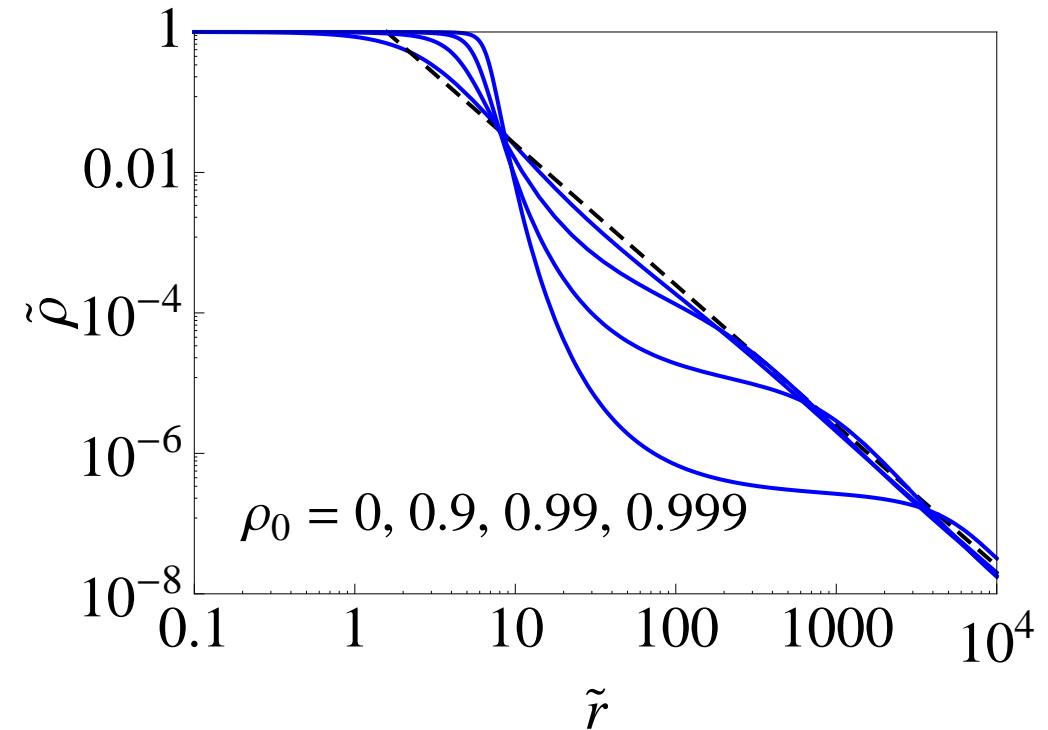


Open ILG Cluster in $\mathcal{D} = 3$



Numerical solution:

- $\tilde{\rho}(\tilde{r})_{\text{as}} \sim \frac{2}{\tilde{r}^2}$
- infinite mass, algebraic decay law
- ICG: Bonnor-Ebert sphere
 - gaseous core
 - gaseous halo
- ILG: emerging horizon
 - condensed matter core
 - gaseous shell
 - gaseous halo

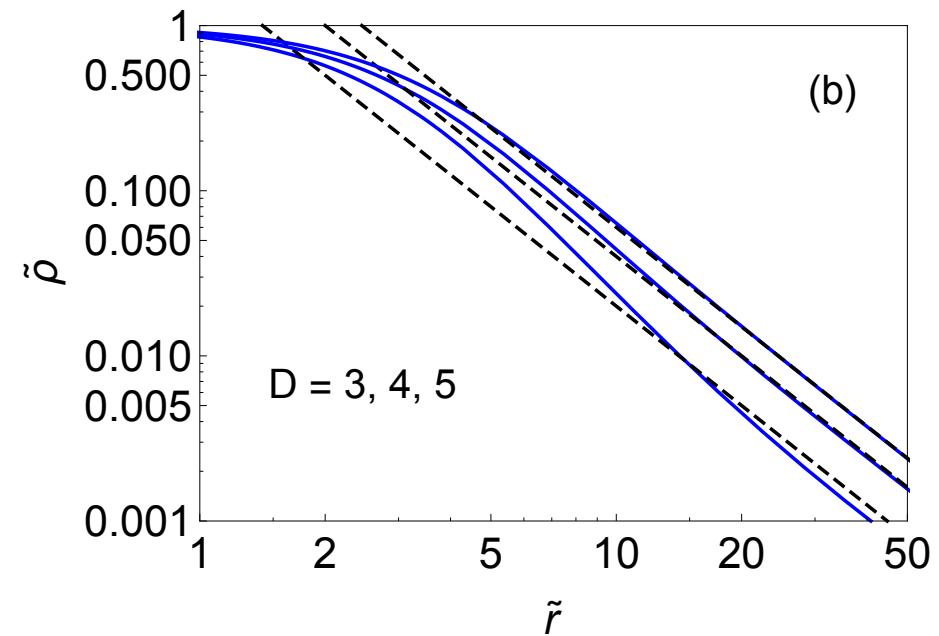
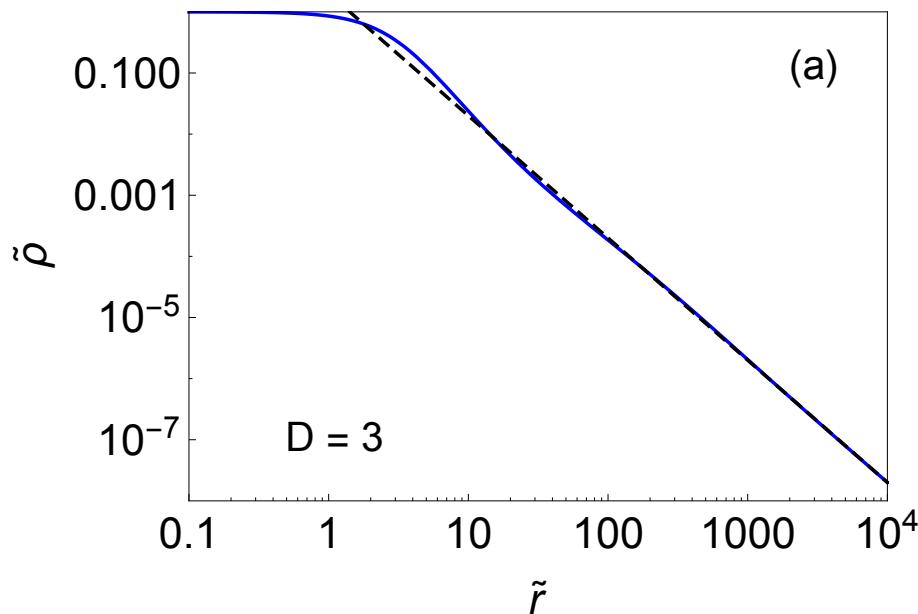


ICG Asymptotics in $\mathcal{D} \geq 3$



$$\frac{\tilde{\rho}''}{\tilde{\rho}} + \frac{\mathcal{D} - 1}{\tilde{r}} \frac{\tilde{\rho}'}{\tilde{\rho}} - \left(\frac{\tilde{\rho}'}{\tilde{\rho}} \right)^2 + \tilde{\rho} = 0 \quad \Rightarrow \quad \tilde{\rho}(\tilde{r}) \sim \begin{cases} e^{-\sqrt{2}\tilde{r}} & : \mathcal{D} = 1, \\ \tilde{r}^{-4} & : \mathcal{D} = 2, \\ \tilde{r}^{-2} & : \mathcal{D} = 3. \end{cases}$$

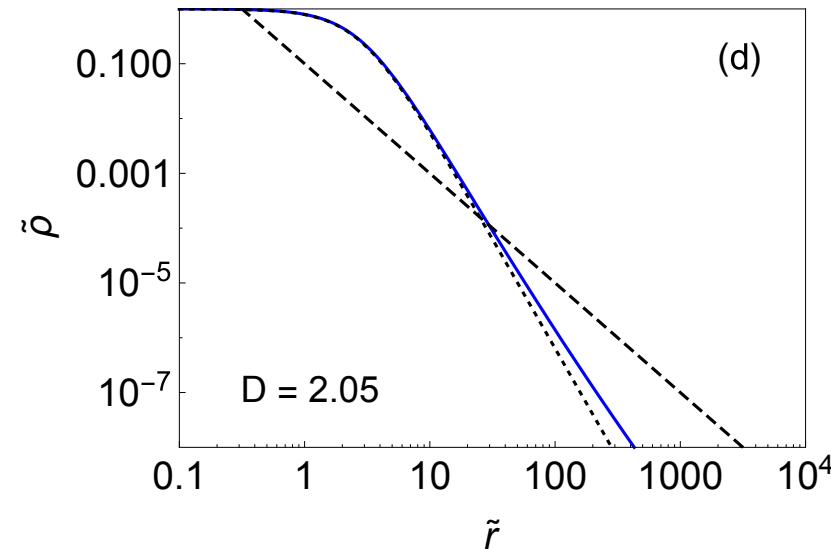
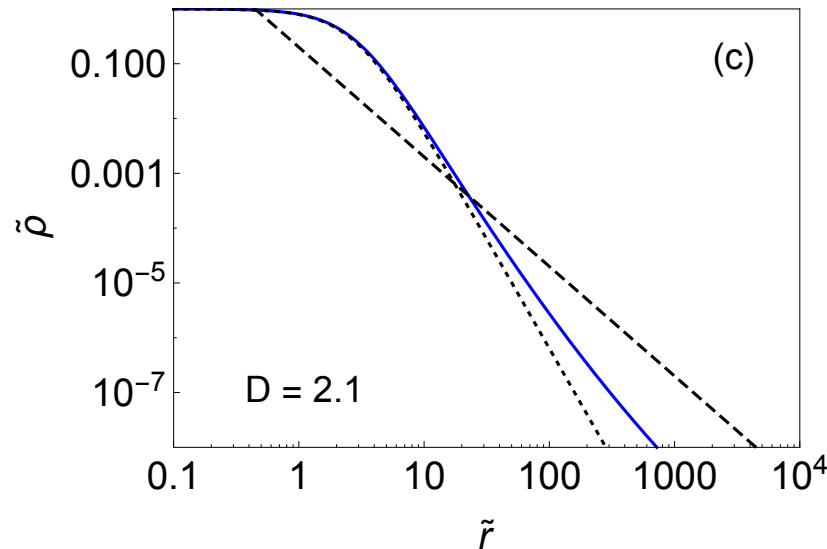
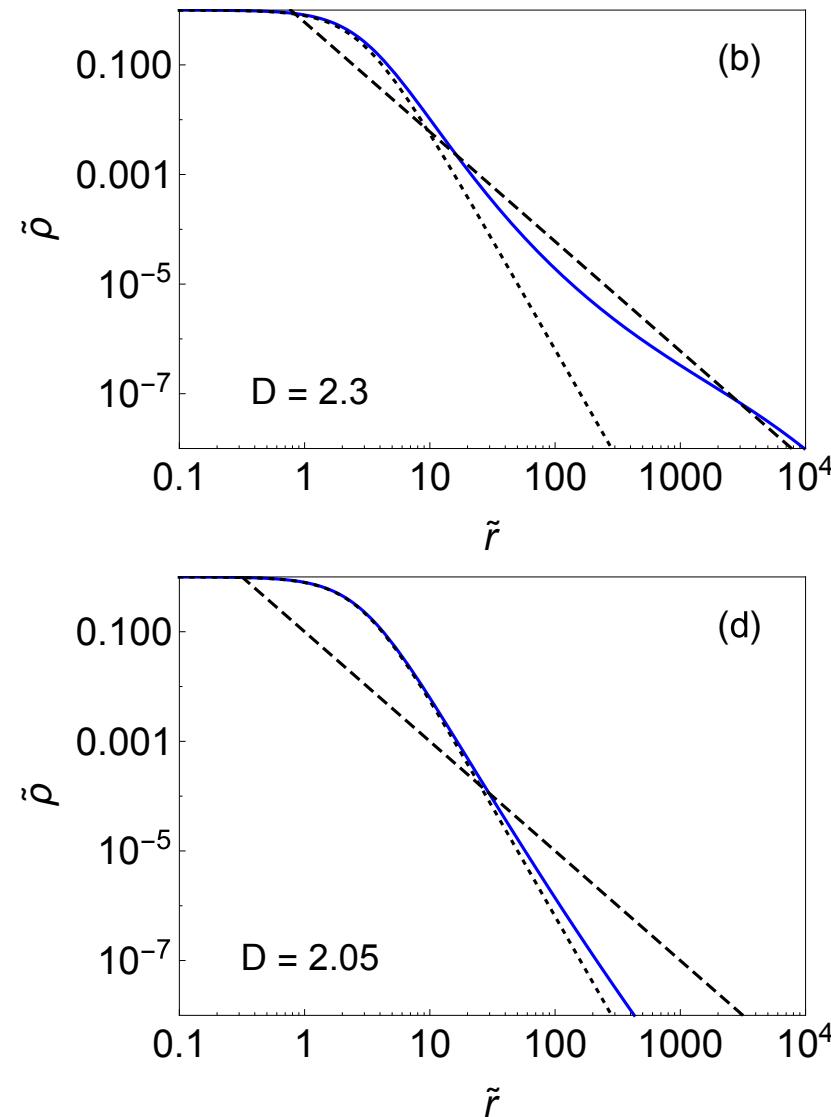
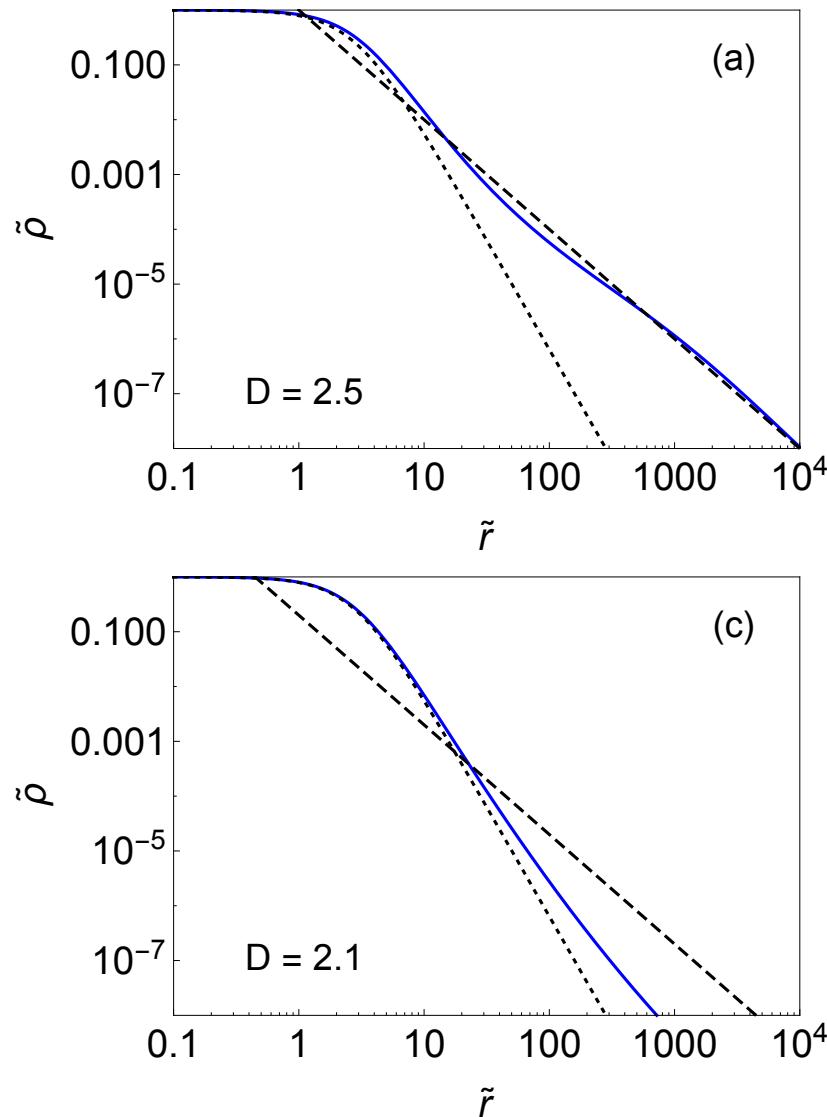
- $\tilde{\rho}(\tilde{r}) \sim a\tilde{r}^{-\alpha}; \quad \alpha = 2, \quad a = 2(\mathcal{D} - 2).$
- infinite mass, algebraic decay law



ICG Asymptotics in $2 \leq D \leq 3$



- infinite mass, crossover between algebraic decay laws

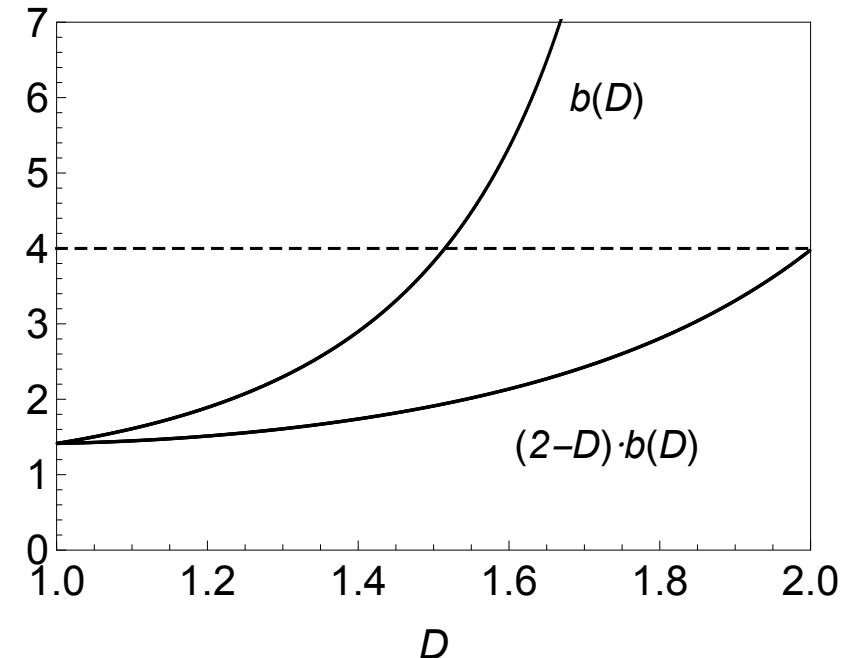
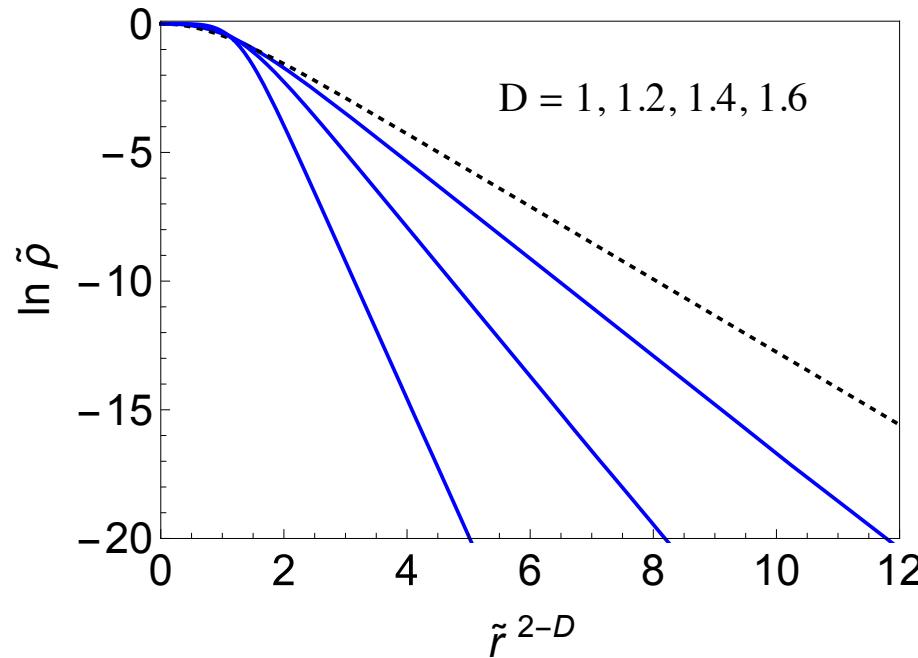


ICG Asymptotics in $1 \leq \mathcal{D} \leq 2$



$$\frac{\tilde{\rho}''}{\tilde{\rho}} + \frac{\mathcal{D}-1}{\tilde{r}} \frac{\tilde{\rho}'}{\tilde{\rho}} - \left(\frac{\tilde{\rho}'}{\tilde{\rho}} \right)^2 + \tilde{\rho} = 0 \quad \Rightarrow \quad \tilde{\rho}(\tilde{r}) \sim \begin{cases} e^{-\sqrt{2}\tilde{r}} & : \mathcal{D} = 1, \\ \tilde{r}^{-4} & : \mathcal{D} = 2, \\ \tilde{r}^{-2} & : \mathcal{D} = 3. \end{cases}$$

- $\tilde{\rho}(\tilde{r}) \sim \exp(-b\tilde{r}^\beta)$; $\beta = 2 - \mathcal{D}$, $\lim_{\mathcal{D} \rightarrow 1} b(\mathcal{D}) = \sqrt{2}$, $\lim_{\mathcal{D} \rightarrow 2} \beta(\mathcal{D})b(\mathcal{D}) = 4$.
- finite mass, stretched exponential decay law





Rotating clusters:

- Competition between gravitational and centrifugal forces.
- Multivalued functional relation between \vec{L} and $\vec{\omega}$.
- instabilities of density profiles:
 - binary configurations in $\mathcal{D} = 1$,
 - ring configurations in $\mathcal{D} = 2$,
 - effects of hysteresis.

Quantum gas clusters:

- ODE for fugacity profile $z(r)$.
- Comparison between FD and BE gases, CS generalization.
- FD Pauli principle vs ILG hardcore repulsion.
- Relativistic effects.
- BE condensation vs gravitational collapse.
- FD cluster drifting into region of magnetic field.